## Fast Light, Fast Neutrinos?

## Kevin Cahill

Department of Physics & Astronomy, University of New Mexico, Albuquerque, New Mexico Physics Department, Fudan University, Shanghai, China\*

(Dated: November 30, 2011)

In certain media, light has been observed with group velocities faster than the speed of light. The recent OPERA report of superluminal 17 GeV neutrinos may describe a similar phenomenon.

Over the past decade, physicists have observed light moving through certain media with group velocities faster than the speed of light in vacuum [1]. The OPERA collaboration [2] may have detected the neutrino analog of this "fast light" phenomenon.

The group velocity of an amplitude

$$A(x,t) = \int e^{i(k \cdot x - \omega t)} B(k) dk$$
 (1)

is determined by the condition that the phase remain constant as a function of wave-number,  $v_g = d\omega/dk$ . When neutrinos traverse a medium with a complex index of refraction n, scattering makes the wave number k complex with a positive imaginary part. The real frequency  $\omega$  is related to the complex wave number k by  $k = n\omega/c$ . If  $n_r$  is the real part of the index of refraction, then the group velocity is

$$v_g = \frac{d\omega}{dk_r} = \frac{c}{n_r + \omega \, dn_r / d\omega}.$$
 (2)

The index of refraction n is related to the forward scattering amplitude f and the density N of scatterers by [3]

$$n(\omega) = 1 + \frac{2\pi c^2}{\omega^2} N f(\omega)$$
 (3)

in which for simplicity I replaced  $k^2$  by  $\omega^2/c^2$ , which introduces an error of less than  $10^{-19}$  for 17 GeV neutrinos with a mass of less than 2 eV/ $c^2$  [2].

Group velocities faster than c can occur when the frequency  $\omega$  is near a resonance in the total cross-section. For instance, if the amplitude for forward scattering is of the Breit-Wigner form

$$f(\omega) = f_0 \frac{\Gamma/2}{\omega_0 - \omega - i\Gamma/2} \tag{4}$$

then the real part of the index of refraction is

$$n_r(\omega) = 1 + \frac{\pi c^2 N f_0 \Gamma(\omega_0 - \omega)}{\omega^2 \left[ (\omega - \omega_0)^2 + \Gamma^2 / 4 \right]}$$
 (5)

and by (2) the group velocity is

$$v_g = c \left[ 1 + \frac{\pi c^2 N f_0 \Gamma \omega_0}{\omega^2} \frac{\left[ (\omega - \omega_0)^2 - \Gamma^2 / 4 \right]}{\left[ (\omega - \omega_0)^2 + \Gamma^2 / 4 \right]^2} \right]^{-1}$$
 (6)

which is superluminal if  $(\omega - \omega_0)^2 < \Gamma^2/4$ .

More generally, we may use the optical theorem and the regularized Kramers-Kronig formula

$$n_r(\omega) = 1 + \frac{cN}{\pi} \int_0^\infty \frac{\sigma_t(\omega') - \sigma_t(\omega)}{\omega'^2 - \omega^2} d\omega'$$
 (7)

in which  $\sigma_t$  is the total cross-section to write the group velocity in terms of the principal part of an integral

$$\frac{c}{v_g(\omega)} = 1 + \frac{cN}{\pi} P \int_0^\infty \frac{\left[\sigma_{\rm t}(\omega') - \sigma_{\rm t}(\omega)\right] (\omega'^2 + \omega^2)}{(\omega'^2 - \omega^2)^2} d\omega'$$
(8)

which shows the effect of scattering on group velocities. Just as the scattering of photons by atoms can cause fast light [1], so too the scattering of neutrinos by electrons and quarks may make neutrino group velocities that are faster than the speed of light. The  $\nu_{\mu}$ -nucleon charged-current total cross-section rises linearly up to 300 GeV [4] and makes a positive contribution to the integral (8). Yet the OPERA Collaboration [2] may have discovered "fast neutrinos"—neutrinos with group velocities faster than the speed of light [5]. Their high group velocity  $(v-c)/c=2.48\times 10^{-5}$  may arise from a resonance in neutrino-electron and/or neutrino-quark scattering at an energy  $\omega_0$  somewhere near 17 GeV.

A group velocity faster than c doesn't violate special relativity, but a superluminal signal velocity would [1].

I should like to thank Franco Giuliani, Lei Ma, and David Waxman for helpful conversations.

- $^*$  cahill@unm.edu
- A. M. Steinberg, P. G. Kwiat, and R. Y. Chiao, Phys. Rev. Lett. 71(5), 708 (1993); Y.-p. Wang and D.-l. Zhang, Phys. Rev. A 52(4), 2597 (1995); M. S. Bigelow, N. N. Lepeshkin, and R. W. Boyd, Science 301(5630), 200 (2003); M. D. Stenner, D. J. Gauthier, and M. A. Neifeld, Nature 425 (2003); N. Brunner, V. Scarani, M. Wegmüller, M. Legré, and G. N., Phys. Rev. Lett. 93(20) (2004).
- [2] OPERA, ArXiv:1109.4897 [hep-ex].
- [3] J. Liu, Phys. Rev. D **45(4)**, 1428 (1992).
- [4] K. Nakamura, J. Phys. G **37**, 075021 (2010).
- [5] V. Panković, ArXiv:1109.6121 [physics.gen-ph].